

TABLE 1  
PROPERTIES OF GALACTIC BLACK HOLE X-RAY BINARIES<sup>a</sup>

No	Name	$M_{\text{comp}}[M_{\odot}]^b$	Spec. type	$M_{\text{BH}}[M_{\odot}]$	$P_{\text{orb}}$	citations <sup>g</sup>
1	XTE J1118+480	$0.22 \pm 0.07$	K7/M1V	$6.9 \div 8.2$	4.08	[1,17,17,38]
2	XTE J1550-564	$0.3 \pm 0.07$	K2/4IV	$10.5 \pm 1.0$	37	[2,2,28,2]
3	GS 2000+25	$0.16 \div 0.47$	K3/6V	$\sim 6.55$	8.26	[3,18,3,40]
4	GRO J0422+32	$\sim 0.45$	M0/4V	$\sim 10.4$	5.09	[4,19,4,4]
5	GRS 1009-45	$\sim 0.5$	G5/K0V	$\sim 8.5$	$6.86 \pm 0.12$	[5,20,5,41]
6	GRS 1716-249	$\sim 1.6$	K-M	$\gtrsim 4.9$	14.7	[37,37,37,37]
7	GX339-4	$0.3 \div 1.1$	KIV	$> 7$	42	[6,21,6,21]
8	H1705-25	$0.15 \div 1.0$	K3/M0V	$4.9 \div 7.9$	12.55	[7,22,29,29]
9	A0620-00	$0.68 \pm 0.18$	K2/7V	$6.6 \pm 0.25$	7.75	[8,23,30,42]
10	XTEJ1650-50(0)	0.7	K4V	$\sim 5.1$	7.63	[ <sup>d</sup> ,24,31,24]
11	XTEJ1859+226	0.7	K5V	$7.7 \pm 1.3$	$6.58 \pm 0.05$	[ <sup>d</sup> ,25,32,25]
12	GS2023+338	$0.5 \div 1.0$	K0/3IV	$12 \pm 2$	156	[9,26,33,43]
13	GRS 1124-68	$0.3 \div 2.5$	K5V	$6.95 \pm 0.6$	10.392	[10,10,34,44]
14	GRS1915+105	$0.8 \pm 0.5$	K1/5III	$12.9 \pm 2.4$	$811.2 \pm 2.4$	[11,27,35,35]
15	GS 1354-64	1.03	G5IV	$7.6 \pm 0.7$	61.07	[ <sup>d</sup> ,12,12,12]
16	GROJ1655-40	$1.75 \pm 0.25$	F3/G0IV	$5.31 \pm 0.07$	$62.909 \pm 0.003$	[ <sup>e</sup> ,13,36,45]
17	4U1543-47	$2.3 \div 2.6$	A2V	$2.7 \div 7.5$	26.8	[14,14,14,46]
18	XTEJ1819-254	$5.49 \div 8.14$	B9III	$8.73 \div 11.70$	67.62	[15,15,38,15]
19	CygX-1 <sup>b</sup>	$19.2 \pm 1.9$	OI	$14.8 \pm 0.1$	134.4	[16,16,16,16]
20	LMC X-1 <sup>c</sup>	$31.79 \pm 3.48$	O7/O8	$10.91 \pm 1.55$	93.84	[47,47,47,47]
21	LMC X-3 <sup>cf</sup>	$3.72 \pm 0.24$	B5III	$7.00 \pm 0.32$	40.8	[48,48,48,48]
22	IC 10 X-1 <sup>c</sup>	$> 17$	WNE	$> 23.1$	$34.93 \pm 0.04$	[49,49,49,49]
23	NGC 300 X-1 <sup>c</sup>	$26^{+7}_{-5}$	WN5	$20 \pm 4$	$32.3 \pm 0.2$	[50,50,50,50]
24	M33 X-7 <sup>c</sup>	$70.0 \pm 6.9$	O7/O8 III	$15.65 \pm 1.45$	82.87	[51,51,51,52]

<sup>a</sup> Most of the data taken from Wiktorowicz et al. (2013)

<sup>b</sup> Cyg X-1 is the only persistent Galactic source, other systems are transients

<sup>c</sup> 5 extragalactic sources. All persistent

<sup>d</sup> Derived from the spectral type

<sup>e</sup> Derived from  $M_{\text{BH}} = 5.31 \pm 0.07$  (Motta et al. 2014) and  $q = 0.329 \pm 0.047$  (González Hernández et al. 2008)

<sup>f</sup> It was suggested by Orosz et al. (2014) that LMC X-3 may be a transient system that is perpetually in the outburst phase (i.e., never drops to the quiescent state with  $L_X < 10^{33} \text{ erg s}^{-1}$ )

<sup>g</sup> Citations are organized as follows: [M<sub>comp</sub>, Spectral type, M<sub>BH</sub>, P<sub>orb</sub>]. References: [1]González Hernández et al. (2012); [2]Orosz et al. (2011b); [3]Ioannou et al. (2004); [4]Reynolds et al. (2007); [5]Macias et al. (2011); [6]Muñoz-Darias et al. (2008); [7]Martin et al. (1995); [8]Gelino et al. (2001b); [9]Casares & Charles (1994); [10]Shahbaz et al. (1997); [11]Harlaftis & Greiner (2004) [12]Casares et al. (2009); [13]Beer & Podsiadlowski (2002); [14]Orosz et al. (1998); [15]Orosz et al. (2001); [16]Orosz et al. (2011a); [17]Khargharia et al. (2013); [18]Harlaftis et al. (1996); [19]Gelino & Harrison (2003); [20]della Valle et al. (1997); [21]Hynes et al. (2003); [22]Filippenko et al. (1997); [23]Froning et al. (2011); [24]Orosz et al. (2004); [25]Corral-Santana et al. (2011); [26]Khargharia et al. (2010); [27]Greiner et al. (2001); [28]Li et al. (2013); [29]Harlaftis et al. (1997); [30]Cantrell et al. (2010); [31]Slaný & Stuchlík (2008); [32]Shaposhnikov et al. (2009); [33]Shahbaz et al. (1994); [34]Gelino et al. (2001a); [35]Hurley et al. (2013); [36]Motta et al. (2014); [37]Masetti et al. (1996); [38]Martin et al. (2008); [39]Torres et al. (2004); [40]Chevalier & Ilovaisky (1990); [41]Shahbaz et al. (1996); [42]Johannsen et al. (2009); [43]Casares et al. (1992); [44]Orosz et al. (1996); [45]González Hernández et al. (2008); [46]Orosz et al. (2002); [47]Orosz et al. (2009); [48]Orosz et al. (2014); [49]Silverman & Filippenko (2008); [50]Crowther et al. (2010); [51]Orosz et al. (2007); [52]Pietsch et al. (2006);

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